



AD-E400 414

TECHNICAL REPORT ARLCD-TR-79010

SAZEL: A COMPUTER PROGRAM TO FIND SUN AZIMUTH AND ELEVATION

EDWARD F. BROWN

APRIL 1980



US ARMY ARMAMENT RESEARCH AND DEVELOPMENT COMMAND
LARGE CALIBER
WEAPON SYSTEMS LABORATORY
DOVER, NEW JERSEY

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REPORT DOCUMENTATION	PAGE	READ INSTRUCTIONS BEFORE COMPLETING FORM
1. REPORT NUMBER	2. GOVT ACCESSION NO.	3. RECIPIENT'S CATALOG NUMBER
Technical Report ARLCD-TR-79010		
4. TITLE (and Subtitle)		S. TYPE OF REPORT & PERIOD COVERED
SAZEL: A COMPUTER PROGRAM TO FIND	SUN AZIMUTH	
AND ELEVATION		
		6. PERFORMING ORG. REPORT NUMBER
7. AUTHOR(*)		8. CONTRACT OR GRANT NUMBER(s)
Edward F. Brown		
9. PERFORMING ORGANIZATION NAME AND ADDRESS		10. PROGRAM FLEMENT PROJECT TASK
ARRADCOM, LCWSL		10. PROGRAM ELEMENT, PROJECT, TASK AREA & WORK UNIT NUMBERS
Applied Sciences Division (DRDAR-	LCA-F)	
Dover, NJ 07801	Jon 1)	
11. CONTROLLING OFFICE NAME AND ADDRESS		12. REPORT DATE
US ARRADCOM		APRIL 1980
ATTN: DRDAR-TSS, STINFO Div		13. NUMBER OF PAGES
Dover, NJ 07801		46 1S. SECURITY CLASS. (of this report)
14. MONITORING AGENCY NAME & ADDRESS(If differen	t from Controlling Office)	13. SECURITY CEMSS. (of this report)
	•	UNCLASSIFIED
		15a, DECLASSIFICATION/DOWNGRADING SCHEDULE
16. DISTRIBUTION STATEMENT (of thie Report)		
Approved for public release; d	istribution unlin	nited.
17. DISTRIBUTION STATEMENT (of the abatract entered	in Block 20, if different fro	m Report)
18. SUPPLEMENTARY NOTES		· · · · · · · · · · · · · · · · · · ·
19. KEY WORDS (Continue on reverse side if necessary at		
Sun position Sun eleva	ation	
Yaw sonde		
Computer program Sun azimuth		
Sun azımutn		·
20. ABSTRACT (Continue en reverse side if necessary an	d identify by block number)	
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A computer program, SAZEL, is presented which computes the sun azimuth and elevation angle from an observer's geographic location, local mean time, and calendar date. The method of computing the sun azimuth and elevation angle is discussed in detail and an application of its use is shown.

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ACKNOWLEDGMENTS

The author is grateful to Mr. Eugene Friedman, a co-worker in the Applied Sciences Division of the Large Caliber Weapons System Laboratory, ARRADCOM, for his suggestions and keen understanding of celestial problems as seen by a navigator. Mr. Friedman was responsible for incorporating SAZEL into the ARRADCOM six degree-of-freedom trajectory program. This resulted in a graphic representation of simulated angular motion from a yaw sonde flight of a free flight missile.

Special thanks are also due to Major L. Rockwell of the U. S. Air Force Reserve, Niagara Falls International Airport, NY. His many hours of instruction and discussion in numerous navigational problems have proven invaluable to the author.

FOREWORD

The azimuth and elevation of the sun are routinely determined by a navigator in the air and on the sea to find his vehicle's direction and line of position (LOP). A sun LOP, when used in conjunction with another LOP obtained from other celestial bodies (moon and Venus) or radio navigational aids, enables the navigator to locate his vehicle's position (latitude and longitude).

The method described in this report to find the sun's azimuth and elevation angles is based on a procedure which the author learned during undergraduate Navigator Training at Mather Air Force Base, CA. It is such an easy method that in a minute's time one can obtain the azimuth and elevation angle of the sun by using two tables: The Air Nautical Almanac and the H.O. 249 Sight Reduction Table.

Rather than manually extracting data from the tables, the necessary celestial mechanics and solar motion data have been programmed into a computer program, SAZEL. One need only know the time of day, position, and date to obtain positional data for the sun.

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INTRODUCTION

To obtain data on rotational motions of projectiles in free flight, a device called a "yaw sonde" is used. The working details of the yaw sonde are described in references 1 through 4. Attitude data are obtained during the flight of a missile by making repeated measurements of the angle between the longitudinal axis of the missile and the sun. Any change in this angle can be approximated by a change in attitude of the missile, since the sun moves infinitesimally during a projectile's flight time.

In order for the yaw sonde to obtain measurements, it is necessary that the sun be within the field of view scanned by the yaw sonde during the flight of the missile. The interval between the first and last times the sun appears in the field of view for the interesting parts of the trajectory on a given day and at a given location is the "window". To determine the window, the sun position (consisting of azimuth and elevation) must be known or computed from other sun data. References 5 through 7 present various computer methods of determining the window. Common to every method, however, is a requirement for selected input data relative to the sun, which is obtained from references 8 and 9.

This report describes various methods of determining the sun's azimuth and elevation (for any year from 1965 through 2000) and, by the use of the computer program SAZEL, eliminates the requirement for input data obtained from references 8 and 9.

DISCUSSION

Sun Features

The sun exhibits certain features that contribute to its use in the yaw sonde method of measuring missile attitude:

- 1. Its position can be determined accurately at any instant in time.
- 2. A change in its position, in relation to the earth, is of so short a duration as to be negligible during the flight of most missiles.
- 3. It provides ample energy for the functioning of the yaw sonde.

Sun Acquisition by a Yaw Sonde

The signals produced by the yaw sonde can be reduced to the spin rate of the projectile and the angle, γ , between the sun vector and the longitudinal axis of the projectile; or the angle, σ , between the sun vector and a normal-to-the projectile axis (fig 2). The convention adopted here is that σ is positive if the sun is toward the nose of the projectile and

$$\sigma = 90^{\circ} - \gamma \tag{1}$$

The ARRADCOM yaw sonde is capable of accepting values of γ from 30° to 145°; then σ may vary from -55° to +60°.

Data reduction of signals for a typical 155 mm M549 projectile in free flight, using a yaw sonde, are then plotted against the time required to study the dynamic behavior throughout its trajectory. Figures 4 and 5 show, respectively, the spin (PHI DOT) and the angular motion (Sigma N) vs time (ref 10).

Time Computation

The Greenwich Mean Time (GMT) of a firing is found by noting the Local Mean Time (LMT) at the instant of firing and converting as follows:

$$GMT = LMT - IZREF (East Longitude)$$
 (2b)

Values of IZREF for each state in the U.S. are given in table 1. For locations outside the U.S., the value of IZREF can be found in reference 8. As an example, the GMT at Yuma Proving Ground, Arizona, for LMT equals 0830 is found to be 1530 since IZREF equals 0700 according to table 1.

Celestial Angle Relations

Figure 1 shows the angular position of the sun and the observer with reference to the earth, along with the symbols and sign conventions. The positional system employed here (ref 12) is one which easily defines the position of an observer on the earth by latitude and longitude. A point on the earth directly beneath the sun, its subpoint (SS, fig 1), could just as easily be defined by its latitude and longitude. The usual point of reference for longitude (fig 1) is the Greenwich Meridian (GM) 0° longitude. The longitude component of the position of the sun's subpoint, SS, is usually expressed as a hour angle either from the Greenwich Meridian (the Greenwich Hour Angle, GHA) or from the observer's

meridan (the Local Hour Angle, LHA). Longitude is measured east or west from 0° through 180° , but hour angles are always measured westward, either from GM or the observer's meridian from 0° through 360° . The Local Hour Angle is computed as follows:

$$LHA = GHA - West Longitude$$
 (3a)

$$LHA = GHA + East Longitude$$
 (3b)

Many angular measurements, such as those shown in figure 1, are referenced to the equator (e.g., the latitude of the observer and the declination (Dec) of the sum) or to the celestial horizon for the elevation of the sum. The declination of the sum is also the latitude of its subpoint. Latitude is measured north or south of the equator, 0° through 90° at the poles. Declination is limited to approximately $23\frac{1}{2}^{\circ}$ travel north or south of the equator (due to the orbit of the earth about the sum and the tilt of the earth's north-south axis).

Finally, we need to define the angular direction in which to look for the sum. Figure 3 shows that the azimuth of the sum, ZN, from an observer's position, is measured clockwise from north to a line joining the observer's position and the subpoint of the sum.

Determination of GHA and Declination of the Sun

Two methods of finding the GHA and Dec of the sun are presented. The first method requires a simple reference to the Air Nautical Almanac for a specific date and GMT. [Table 2 is extracted from the almanac (ref 8).] As an example, for 28 Oct 1978 at 1530 GMT, the almanac states that the GHA equals 56° 32.5' (56.542°) and the Dec = S 13° 08.4' (-13.14°). The second method (ref 11) employs the use of table 3 and is explained in the table. It is an emergency method of computing the sun's GHA and Dec if the almanac is not available. At first glance, it seems cumbersome; but, with practice, the computation is accomplished in a minute's time. Using the same example as above and following the explanation in table 3, we would compute GHA equals 56° 33' (56.55°) and Dec equals S 13° 09' (-13.15°). This method has been programmed into the computer program, SAZEL. User input to the program is explained in appendix A and a listing of SAZEL is provided in appendix B.

Determination of the Sun's LHA, Elevation, and Azimuth

Once the GHA and Dec of the sun have been found, based on the date and GMT, the azimuth (ZN) and elevation (HC) of the sun can be found if we know the latitude and longitude of the observer. Through the use of the Sight Reduction Tables for Air Navigation (ref 11),

we can find the azimuth and elevation of the sun. (Table 4 is an extract from reference 11 to use as an example.) Using the GHA and Dec found earlier, and placing the observer at Yuma Proving Ground, latitude equals 32.8811° and longitude equals 114.3028°, we proceed to find the LHA using equation 3a, thus

LHA = $56.5500^{\circ} - 114.3028^{\circ}$

LHA = 302,2472° (360° has been added to make LHA positive)

Entering table 4 with the LHA, Dec, and latitude, we find the azimuth and elevation of the sun by interpolation; thus, HO = 18° 11' and Z = ZN = 120° . The computer program, SAZEL, computes HC and ZN using rudimentary principles of trigonometry.

Six Degrees-of-Freedom Simulation of Free Flight Missile Using SAZEL

For a simulated trajectory of a free flight missile, using a modified six degrees-of-freedom (6-DOF) trajectory program (ref 13) and SAZEL, the angular motion, σ , can be plotted against time as shown in figure 6. Figure 6 represents a simulation of the 155 mm M549 projectile which was fired at Yuma Proving Ground (ref 10) compared to the actual recorded angular motion shown in figure 4. The azimuth and elevation of the sun can be represented by a vector S, as shown in figure 3. The vector S can further be represented by unit vector components S1, S2, and S3. The attitude of the missile, as contained in reference 13, is represented by the vector P. Vector P is also represented by its unit vector components P1, P2, and P3. The dot product of these unit vectors enables us to find γ and finally, σ , which we are ultimately interested in plotting against time. S1, S2, and S3 in the earth frame are as follows:

$$S1 = \cos HC \cos ZN$$
 (4a)

$$S2 = \sin HC \tag{4b}$$

$$S3 = \cos HC \sin ZN$$
 (4c)

The angle γ is then computed as follows:

 $\gamma = arc \cos (51P1 + S2P2 + S3P3)$

and ultimately we find σ from equation 1.

Accuracy

The calibration of the yaw sonde is such that system acquisition has been reported to be $\pm \frac{1}{4}$ °. The method of computation using SAZEL is reported in the Air Nautical Almanac to be ± 1' (0.017°) for GHA and $\pm 2'$ (0.033°) for Dec of the sun. Small errors in the computation of azimuth and elevation would be inherent due to GHA and Dec from the above computation and changing missile position. The effect of changing latitude and longitude after launch on the computation of GHA, LHA, HC, and ZN, can be minimized by using the latitude and longitude at the midpoint of a trajectory. Another error is introduced through the necessary assumption of unchanging altitude from ground level. Finally, an error which affects both computation of sun elevation and data acquisition is atmospheric refraction. Refraction is a complicated function of height above the earth and apparent elevation of the sun. fraction is further dependent on the temperature of the atmosphere at the height of observation. At any given apparent sun elevation, the refraction decreases with increasing altitude. Refraction is further decreased if the atmosphere is warmer than normal at the height of observation. For the flight of most missiles, the refraction encountered should be no greater than -20' (0.333°) under the worst conditions-sea level, colder than normal atmosphere, and sun on the horizon.

The overall error for ranges under 30km and altitudes from ground level to 30 km, using the midpoint trajectory latitude and longitude, is estimated to be $\pm \frac{120}{2}$ for any simulation.

RESULTS

The reader can compare the manual method of computation for LHA, GHA, AZ and HC to that derived from SAZEL. Very close agreement is shown between the results obtained from the SAZEL program (app A) and those extracted from the Air Almanac and Sightreading Tables, as follows:

chose extracted from	CHC ALL ALMA	iac and big	ncreaurng i	autes, as to	IIOWS:
Source	LHA (computed)	GHA	DEC	AZ	HC
Air Almanac		56.5417°	-13,1400°		
Sightreading Tables	302.2472°	56.5500°	-13.1500°		
Sight Reduction Method	302.2472°			120.0000°	18.1833°
SAZEL		56.5333°	-13.1392°	119.8569°	18.2251°

Comparing the actual angular motion in figure 4 to the simulated angular motion in figure 6, close agreement is obvious. Note that the actual flight would result in Sigma N from the apparent sun due to refraction. A suitable aerodynamic coefficient package and the test data (record of firing data for the actual flight) are necessary to obtain accurate results.

CONCLUSIONS

SAZEL, a computer program, has been developed which determines the azimuth and elevation angles of the sun quickly and accurately, without using tabulated sun data available in the published literature. In comparison with the yaw-sonde method (which does rely on previously published data), SAZEL has been demonstrated to be a viable tool in simulating the angular motion of missiles during free flight.

RECOMMENDATIONS

- 1. Update, with SAZEL, the ARRADCOM computer program for determining appropriate times (windows) for firing projectiles equipped with yaw sondes. (The trial version is now working.)
- 2. Complete the 6-DOF trajectory program using SAZEL which allows computation of projectile angular motion motion (yaw sonde sun angle, Sigma N) vs time for plotting on the ARRADCOM graphics terminal. (The trial version is now working.)

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Table 1. IZREF conversion from local time to Greenwich time

STATE	IZREF
	LZKEF
Alabama	0600
Alaska, east of long. W. 137°	0800
long. W. 137° to W. 141°	0900
long. W. 141° to W. 161°	1000
long. W. 161 ⁰ to W. 172 ⁰ Aleutian Islands	1100
Arizona	1100
Arkansas ¹	0700
California ¹	0600 0800
Colorado ¹	0700
Connecticut ¹	0500
Delaware ¹	0500
District of Columbia ¹ Florida ^{1,2}	0500
Florida ¹ , ²	0500
Georgia ¹	0500
Hawaii	1000
Idahol, 2	0700
Illinois ¹	0600
Indiana ²	0500
Iowa ¹ ,2 Kansas ¹ ,2	0600
Kentucky ¹ , ²	0600
Louisianal	0500
Mainel	0600 0500
Maryland ¹	0500
Massachusetts ¹	0500
Michigan	0500
Minnesota ¹	0600
Mississippi ¹	0600
Missouri ¹	0600
Montana ¹	0700
Nebraska, 2	0600
Nevada ¹	0800
New Hampshire ¹	0500
New Jersey ¹	0500
New Mexico ¹	0700
New York ¹	0500
North Carolina ¹ North Dakota ^{1,2}	0500
Ohio ¹	0600
0klahoma,1	0500 0600
Oregon, 2	0800
Pennsylvania 1	0500
Rhode Island ¹	0500
South Carolina ¹	0500
South Dakota, eastern part	0600
western part	0700
Tennessee 1,2	0600
Texas ¹	0600
Utah, 2	0700
Vermont ¹	0500
Virginia ¹	0500
Washington, D.C. ¹ Washington ¹	0500
Washington 1	0300
West Virginia ¹	0500
"ISCONSIN	0600
Wyoming	0700

Summer (daylight-saving) time, one hour fast on the time given (-100) is kept in these states from the last Sunday in April to the last Sunday in October, changing at $02^{\rm h}$ $00^{\rm m}$ local clock time.

 $^{^2{\,}^{\}phantom{\dagger}}_{}$ This applies to the greater portion of the state.

Table 2. Extract from Air Nautical Almanac

602 (D/AY/301) GREENWICH P. M. 1978 OCTOBER 28 (SATURDAY)

602	(DA1930	I) GRE	ENWICH P.	M. 1978 C	OCTOBER 2	8 (SATURD	AY)
GMT	⊙ \$UN GHA Dec.	ARIES GHA 57	VENUS-4.0 GHA Dec.	JUPITER-1.7 GHA Dec.	SATURN 1.1 GHA Dec.	MOON GHA Dec.	Lat. Moon- set
12 00 10 20 30 40	6 32.3 05.7 9 02.3 05.8 11 32.3 · 05.9 14 02.4 06.1 16 32.4 06.2	218 59.0 221 29.5 223 59.9 226 3C.3 229 0C.7	350 05 S24 04 352 35 355 06 357 36 · · · 0 07 2 37	86 11 N18 41 88 41 91 12 93 42 · · 96 12 98 43	53 18 N 8 52 55 48 58 19 60 49 · · 63 19 65 50	41 48 N 2 05 44 13 03 46 38 02 49 04 2 00 51 29 1 59 53 55 57	72 15 37 -02
13 00 10 20 30 40 50	24 02.4 06.6 26 32.4 06.8 29 02.4 06.9 31 32.4 07.1	234 01.5 236 31.9 239 02.3 241 32.7 244 03.2	5 08 S24 04 7 38 10 09 12 40 · · 15 10 17 41	101 13 N18 41 103 43 106 14 108 44 · · 111 15 113 45	68 20 N 8 52 70 51 73 21 75 51 · · 78 22 80 52	56 20 N 1 55 58 45 54 61 11 52 63 36 50 66 02 49 68 27 47	62 15 30 07 60 15 29 09
10 20 30 40 50	44 02.5 07.7 46 32.5 07.9	249 04 0 251 34 4 254 04 8 256 35 2 259 05 6	20 11 S24 03 22 42 25 12 27 43 · · · 30 14 32 44	116 15 N18 41 118 46 121 16 123 46 · · 126 17 128 47	83 23 N 8 52 85 53 88 23 90 54 · · 93 24 95 54	70 52 N 1 45 73 18 44 75 43 42 78 09 40 80 34 39 83 00 37	50 15 26 13 45 15 25 15 40 15 24 16 35 15 23 17 30 15 22 18 20 15 21 20
30 30 40 50	54 02.5 08.3 56 32.5 08.4 59 02.5 08.6 61 32.5 08.7	264 06 4 266 36 8 269 07 3 271 37 7 274 08 1	35 15 \$24 03 37 45 40 16 42 47 · · 45 17 47 48	131 18 N18 41 133 48 136 18 138 49 · · 141 19 143 49	98 25 N 8 52 100 55 103 25 105 56 · · 108 26 110 57	85 25 N 1 36 87 50 34 90 16 32 92 41 31 95 07 29 97 32 27	10 15 20 22 0 15 18 23 10 15 17 25 20 15 16 26 30 15 15 28
16 00 10 20 30 40 50	69 02.5 09.1 71 32.5 · 09.3	276 38.5 279 08.9 281 39.3 284 09.7 286 40.1 289 10.5	50 18 S24 02 52 49 55 19 57 50 60 21 62 51	146 20 N18 41 148 50 151 21 153 51 · · · 156 21 158 52	113 27 N 8 52 115 57 118 28 120 58 · · · 123 28 125 59	99 57 N 1 26 102 23 24 104 48 22 107 14 21 109 39 19 112 04 17	35 15 14 29 40 15 13 30 45 15 12 31 50 15 10 33 52 15 10 34 54 15 09 35
17 00 10 20 30 40 50	84 02.6 10.0 86 32.6 10.1 89 02.6 10.3	291 41.3 294 11.4 296 41.3 299 12.2 301 42.5 304 13.3	65 22 524 02 67 52 70 23 72 53 · · · 75 24 77 55	161 22 N18 41 163 52 166 23 168 53 171 24 173 54	128 29 N 8 52 131 00 133 30 136 00 · · 138 31 141 01	114 30 N 1 16 116 55 14 119 20 12 121 46 11 124 11 09 126 37 08	56 15 08 35 58 15 08 37 60 15 07 38 S
	101 32.6 · 11.0 104 02.6 · 11.1	306 43.4 309 13.3 311 44.2 314 14.7 316 45 319 15.3	80 25 S24 01 82 56 85 26 87 57 · · 90 27 92 58	176 24 N18 41 178 55 181 25 183 55 · · 186 26 188 56	143 32 N 8 52 146 02 148 32 151 03 · · 153 33 156 03	129 02 N 1 06 131 27 04 133 53 03 136 18 1 01 138 44 0 59 141 09 58	Moon's P. in A.
19 00 10 20 30 40 50	111 32.7 114 02.7 116 32.7 11.8 119 02.7	321 45.9 324 16.3 326 46.7 329 17.1 331 47.5 334 17.4	95 29 S24 00 97 59 100 30 103 00 · · 105 31 108 02	191 27 N18 41 193 57 196 27 198 58 · · · 201 28 203 58		143 34 N 0 56 146 00 54 148 25 53 150 51 51 153 16 49 155 41 48	7 56 56 31 13 55 58 30 16 54 59 29 20 53 60 28 22 52 61 27 25 51 62 26
20	129 02.7 131 32.7 134 02.7 12.6 12.8	339 18.8 341 49.2 344 19.6 346 50.0	110 32 524 00 113 03 115 33 118 04 · · · · · · · · · · · · · · · · · ·	208 59 211 30 214 00 · · 216 30	176 06 178 37 181 07 183 37	158 07 N 0 46 160 32 44 162 57 43 165 23 41 167 48 39 170 14 38	27 50 64 25 29 49 65 24 31 48 66 23 33 47 67 22 35 46 68 21 37 45 68 20
20 30 40 50	144 02.8 13.3 146 32.8 13.5 149 02.8 13.6 151 32.8 13.7	351 50.E 354 21.Z 356 51.6 359 22.C 1 52.5 4 22.9	128 06 130 37 133 07 · · 135 38	224 01 226 32 229 02 · · 231 33	191 09 193 39 196 09 · · · 198 40	172 39 N 0 36 175 04 34 177 30 33 179 55 31 182 20 29 184 46 28	38 44 70 19 40 42 71 17 42 41 73 16 43 40 74 15
10 20 30 40	154 02.8 S13 13.9 156 32.8 14.0 159 02.8 14.2 161 32.8 14.3 164 02.8 14.4 166 32.8 14.6	9 23.7 11 54.1	143 10 145 40 148 11 · · · 150 41	239 04 241 34 244 04 · · · 246 35	208 41 211 12 · · 213 42	187 11 N 0 26 189 37 24 192 02 23 194 27 21 196 53 19 199 18 18	47 38 77 13 49 37 78 12 50 36 79 11 51 35 80 10 53 34 53 34
10 20 30 40 50	169 02.8 S13 14.7 171 32.9 14.9 174 02.9 15.0 176 32.9 15.1 179 02.9 15.3 181 32.9 15.4	24 26.1 26 56.6	158 13 160 44 163 14 165 45	254 06 256 36 259 07 · · · 261 37	221 13 223 43 226 14 · · · 228 44	201 43 N 0 16 204 09 14 206 34 13 208 59 11 211 25 09 213 50 08	54 32 55 32 56 31 Sun SD 16'1 Moon SD 15'
_ Rate	15 00.0 S0 00.8		15 03.5 NO 00.5	15 02.2 S0 00.1	15 02.3 S0 CO.1	14 32.3 50 09.9	Age 26d

GHA and declination of the sun for the years 1965-2000 Table 3.

(a) Interpolation for E and Dec

Ρď	90	8 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	
DEC	Dec.	• 12 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	
a	ш	0 1 4 4 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8	
NOV	Dec.	. 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	
Z	ш	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	
OCT	Dec.	S 2 25.7 S 2 25.7 S 2 25.7 S 3 2 25.7 S 4 2 3 3 4 2 3 4	
	ш	7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7	0. T
SEP	Dec.	N N N N N N N N N N N N N N N N N N N	of
	ш	7 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	Hours
AUG	Dec.	N N N N N N N N N N N N N N N N N N N	on for
.1	ш	0 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	lati
JUL	Dec.	X X X X X X X X X X X X X X X X X X X	Interpolation
	ш	### ### ### ### ### ### ### ### #### ####	(c) I
NOS	Dec.	X X X X X X X X X X X X X X X X X X X	
	ш	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	2
MAY	Dec.	N 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	
_	ш	2 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	
APR	Dec.	A 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	
7	ш	5 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	
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After Feb. 29

i		

(Cont'd) Table 3.

(e) Minutes and Seconds of GMT

(d) Hours and Tens of Minutes of GMT

200	4 00	10m	20m	30m	40m	50m
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_	_		_			
	_		_			
	_		_			
3	235 00		_			
5						
			_		_	
-						292 30
8					_	
8	310 00	312 30	315 00	317 30	320 00	
_						
7						
				17 30	20 00	22 30
14	25 00	27 30	30 00	32 30		37 30
9			_		20 00	
9			_		_	
7			_	77 30	80 00	82 30
8			_		_	
61	100 001	102 30	105 00	107 30	_	112 30
0					_	
			_		_	
22	145 00 3	147 30	150 00	152 30	155 00	157 30

€ 8

EXPLANATION

In leap years the upper value of the h correction 0.T.'s corresponding to GMT's of 1631 February 29 time during the years 1965-2000. Table 3a given E(5 + Equation of Time) and declination of the Sun for the argument "Orbit Time" O.T., the latter formed by applying the h correction from and 0529 March 1, 1968, are 0400 February 29 and is to be used for January and February, and the Thus, of the GHA and declination of the Sun for any Table 3b to the nearest integral hour of GMT. These data make possible the determination lower h value for the rest of the year. 1600 March 1, respectively.

with number of hours of 0.T. as vertical argument. determined by entering Table 3c with differences between consecutive values of E and declination as horizontal argument, and the vertical column Corrections to E and declination for O.T. are

2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	
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5

The GHA is obtained by adding to the corrected E the value of the diurnal arc obtained from Tables 3d and 3c. The latter two rables must be entered with argument GMT.

9°02'(+01') S 12°55'(+20') O.T.=GMT(nearest integral hour)+Corr.(Table 3b) s 13^o09 +14 s 13°09 EXAMPLE - October 28, 1978, $\text{GMT=}15^{\text{h}}30^{\text{m}}01^{\text{s}}$ 0.T.=Oct. $28^{d}16^{h} + 01^{h} = 0ct. 28^{d}17^{h}$ 9003 56033 47030 +01 00 Table 3a, Oct. 28^dO.T. GMT Table 3d, 15h30m GMT E, Dec., Oct. 28^d17^h Table 3c, 17^h0.T. Table 3e, 0^m01^s Sum, GHA, Sun Sum, Dec. Sun

Table 4. Extracts from sight reduction tables

	Zn=180-2 Zn=180+Z			L	AT 32°		V				L	AT 33	•	11	
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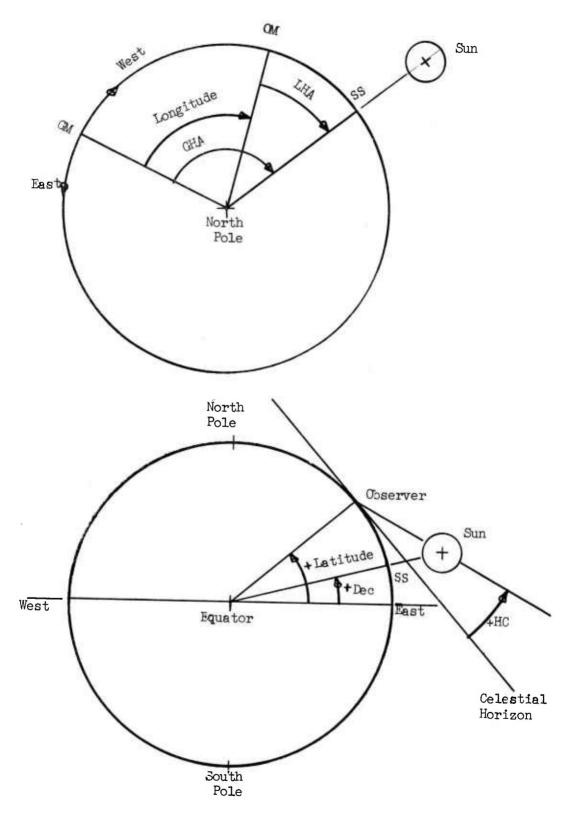


Figure 1. Celestial angle relations.

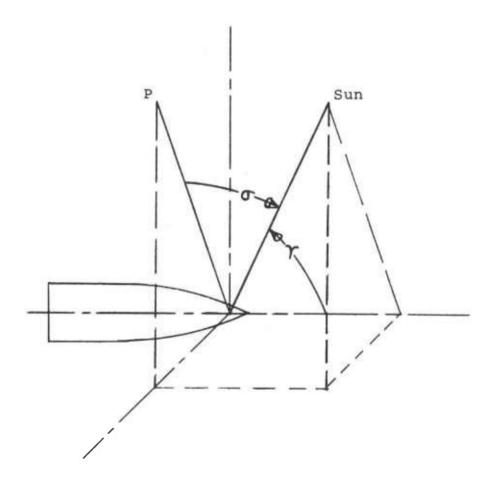


Figure 2. Definition of gamma and sigma.

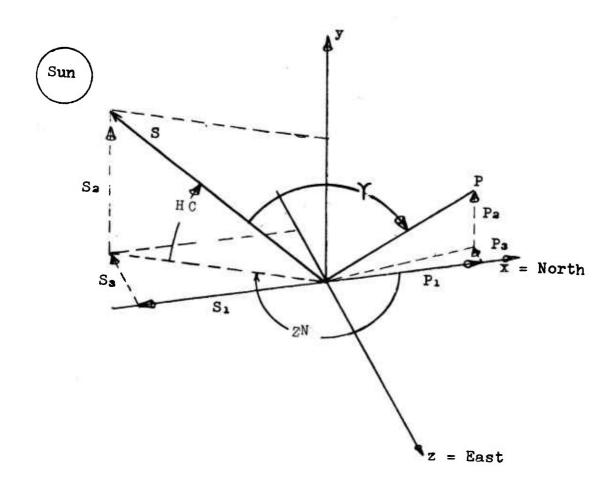
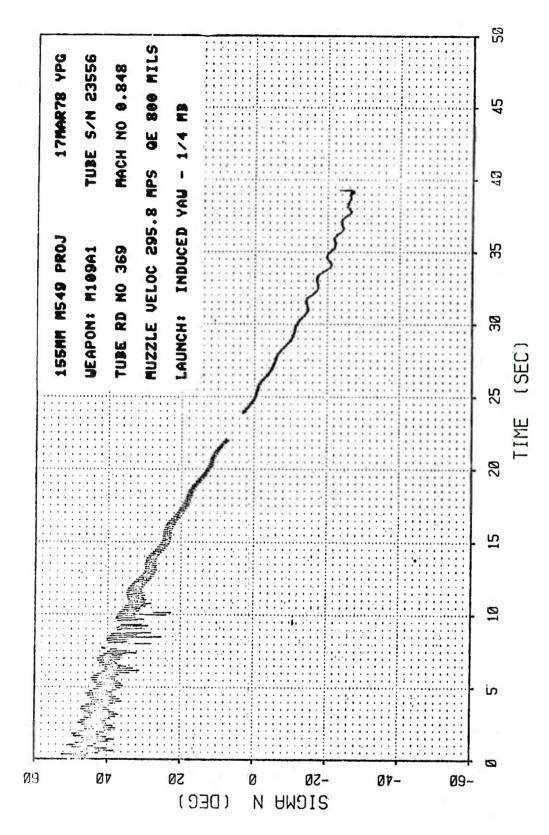
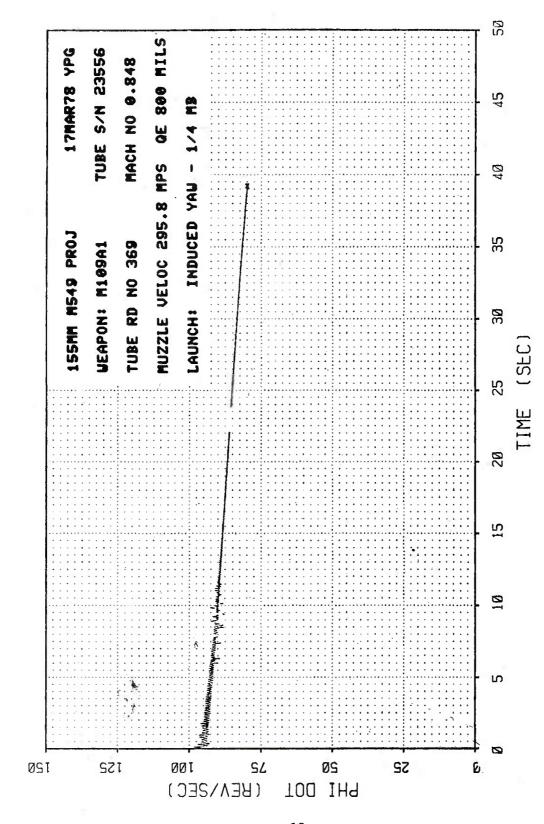


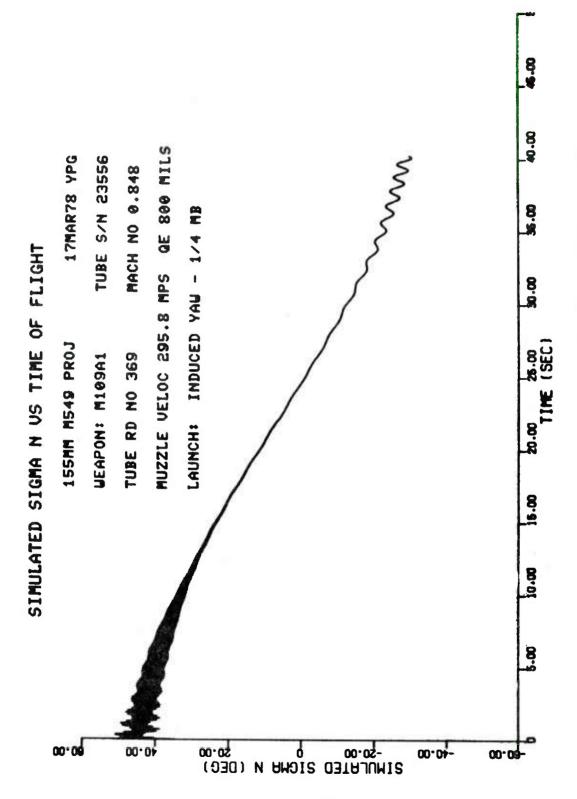
Figure 3. Sun-to-projectile angle resolution and sun s elevation and azimuth.



Actual sigma N vs time of a 155 mm M549 projectile. Figure 4.



Actual phi dot ys time of a 155 mm M549 projectile. Figure 5.



Simulated sigma N vs time of a 155 mm M549 projectile. Figure 6.

APPENDIX A

INPUT AND OUTPUT FOR COMPUTER PROGRAM "SAZEL"

Description of Input and Output of SAZEL

To find the azimuth and elevation of the sun using the computer program, SAZEL, as listed in Appendix B, data are input on one computer card as follows:

INPUT

Column No.	Item	Format
1-2	NDAY, numerical day of month	12
10-11	NMONTH, numerical month of year	12
20-23	NYEAR, numerical year A.D. (1978-2000 only)	14
30-31	NHOUR, GMT hours	12
32-33	NMIN, minutes	12
50-59	WLONG, longitude, degrees	F9.4
60-69	WLAT, latitude, degrees	F9.4

EXAMPLE

The example used in table 3 and 4 is chosen for comparative purposes.

October 28, 1978; GMT=1530 Yuma Proving Ground; Longitude=114.3028 deg, Latitude=32.8811 deg

Column No				Card	l Item		
1-2				28			
10-11				10			
20-23				1978	3		
30-31				15			
32-33				30			
50-59				114.	3028		
60-69				32.8	8811		
OUTPUT							
ODAY	OTIME	(Z)	GHA		AZ IMUTH	DECLINATION	ELEVATION
302	1600		56,533	33	119.8569	-13,1392	18,2251

APPENDIX B

LISTING OF COMPUTER PROGRAM "SAZEL"

1 C CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC	
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C DATEN=JULIAN DATE C GHAS=GREENWICH HOU C LONGITUGE, OE C OECS=SUN OECLINATI C OECS=SUN OECLINATI C TIME JULIAN SUN II C AZHBO=AZIMUTH OF SU C AZHBO=AZIMUTH OF SU C AZHBO=AZIMUTH OF SUN II C AZHBO-E HORIZON C AZHBO-E HORIZON C AZHBO-E HORIZON C BIMENSION MONTH (1) OIMENSION OECI(186 OIMENSION OECI(186)	
C GHAS=GREENING INFOLD C C OECS=SUN OECLINATI C C TIME=JULIAN SUN TI C AZHBO=AZIMUTH OF S C C C C C C C C C C C C C C C C C C C	
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C TIME=JULIAN SUN TIME TO COMPUTE SUN COORDINATES. MAY C AZHBO=AZIMUTH OF SUN FROM TRUE NORTH MEASURED CLOCKWISE C AZHBO=AZIMUTH OF SUN FROM TRUE NORTH MEASURED CLOCKWISE C FROM OBSERVERS POSITION. DEGREES C HBELEVATION OF SUN MEASURED FROM OBSERVERS GROUND C LEVEL HORIZON C EVEL HORIZON C BELOW HORIZON C BELOW HORIZON C DIMENSION DECICLOSCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC	
C AZHBO=AZIMUTH OF SUN FROM TRUE NORTH MEASURED CLOCKWISE C AZHBO=AZIMUTH OF SUN FROM TRUE NORTH MEASURED CLOCKWISE C FROM OBSERVERS POSITION, DEGREES C HB=ELEVATION OF SUN MEASURED FROM OBSERVERS GROUND C LEVEL HORIZON C - BELOW HORIZON C - BELOW HORIZON C C CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC	
C AZHBO=AZIMUTH OF SUN FROM TRUE NORTH MEASURED CLOCKWISE FROM OBSERVERS POSITION, DEGREES C HB=ELEVATION OF SUN MEASURED FROM OBSERVERS GROUND C LEVEL HORIZON C + ABOVE HORIZON C - BELOW HORIZON C - BELOW HORIZON C C CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC	
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	AND DEC2 CON	
09	C AS A FUNCTION OF JULIAN DATE	
	C OECAT AND OECAZ CONTAIN THE OECLINATION INCREMENT FROM	The state of the s
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	\$2,4333;2,5,2,5667;2,6333;2,7,2,7667;2,8333;2,9167;2,9833;3,0667;	
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	120	CONTINUE				
400	4	DATE INDUTHINANTH TYPARI - NDAY	NAV.			
522		IF (IYEAR.EG.2.AND.DATEJ.GT.60) N=((MYEAR-12)/4)-17	60) N= ((MYEAR+12)/4)+17			ľ
		ADD=CORP (N) *100.				
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230		.01 DA				
			D.CH.EG.2400.) DATEN=DATEN+1			
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GLOSSARY

Azimuth

The angle measured clockwise from true north to a line passing through the observer and subpoint of the sun

Celestial horizon

The tangent plane to the earth passing through the observer's position

Declination

The angular distance to the sun measured north or south through 90° from the equator to the subpoint of the sun

Elevation

The angular distance of the sun above the celestial horizon

Greenwich Hour Angle

The angular distance measured from the Greenwich meridian westward through 360° to the meridian passing through the subpoint of the sun

Greenwich Meridian

The prime meridian which passes through Greenwich, England, from which longitude is measured east or west

Greenwich Mean Time

Local time at the Greenwich meridian measured by reference to the mean sun. It is the angle measured along the equator (and converted to time) from the Greenwich meridian westward through 360° to the meridian passing through the subpoint of the mean sun

IZREF

A time conversion to express local mean time as Greenwich mean time

Latitude

Angular distance measured north or south of the equator along a meridian, 0° through 90°

Longitude

The angular distance east or west of the Greenwich meridian, measured along a line of parallel from 0° to 180°

Local Hour Angle

The angular distance measured from the observer's meridian westward through the subpoint of the sun Local Mean Time

Local time at the observer's meridian measured by reference to the mean sun. It is the angle measured along the equator (and converted to time) from the observer's meridian westward through 360° to the meridan passing through the subpoint of the mean sun

Meridian

Imaginary line on the earth connecting points of equal longitude

Paralle1

Imaginary line on the earth connecting points of equal latitude

Subpoint

That point on the earth's surface directly beneath a celestial body

Time

Usually expressed in four numerals (0001 thru 2400) where there are 24 hours in a day and 60 minutes to one hour. 6:29 PM would be expressed as 1829

Z or Zulu time

An expression indicating Greenwich mean time

LIST OF SYMBOLS

Dec Declination of the sun, deg

DOF Degrees-of-freedom

GHA Greenwich Hour Angle of the sun, deg

GM Greenwich Meridian, 0° longitude

GMT Greenwich Mean Time, hours

HC Elevation of the sun, deg

IZREF Time conversion, hours

LHA Local Hour Angle of the sun, deg

LMT Local Mean Time, hours

OM Observer's meridian, deg of longitude

P₁, P₂, P₃ Directional cosines of missile position, P, as used in 6-DOF trajectory program (earth frame)

S₁, S₂, S₃ Directional cosines of sun position, S, (earth frame)

SS Sun's subpoint meridian, deg

ZN True azimuth of the sun, deg

Yaw sonde acquisition angle between the sun and pro-

jectile axis, deg

Angle between the sun and transverse plane perpendicular to projectile longitudinal axis, deg

SIGN CONVENTION

Dec North + South -

HC Above celestial horizon +

Below celestial horizon -

UZREF East longitude position -

West longitude position +

Latitude North +

South -

Longitude East +

West -

Forward of projectile + Aft of projectile -Sun

ZN Clockwise +

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